# New asymptotic expansion for the $\Gamma(z)$ function. 

Gergő Nemes<br>Institute of Mathematics, Eötvös Loránd University 1117 Budapest, Hungary

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#### Abstract

Using a functional substitution and Faà di Bruno formula, Stirling asymptotic series approximation to the Gamma function is converted into a new one with better convergence properties. The new formula is compared with those of Stirling, Laplace and Ramanujan for real arguments greater than 0.5 and turns out to be, for equal number of 'correction' terms, numerically superior to all of them. As a side benefit, a closed-form approximation has turned up during the analysis which is about as good as 3rd order Stirling's (maximum relative error smaller than $1 \mathrm{e}-10$ for real arguments greater or equal to 10 .)


1. The Bell polynomials. Suppose that $h(z)=f(g(z))$ and let

$$
h_{n}=\frac{d^{n} h(z)}{d z^{n}}, f_{n}=\frac{d^{n} f(z)}{d z^{n}}, g_{n}=\frac{d^{n} g(z)}{d z^{n}}
$$

then by applying the chain rule the first few derivatives of $h(z)$ are given by

$$
\begin{gathered}
h_{1}=f_{1} g_{1}, \\
h_{2}=f_{1} g_{2}+f_{2} g_{1}^{2}, \\
h_{3}=f_{1} g_{3}+f_{2} 3 g_{1} g_{2}+f_{3} g_{1}^{3} .
\end{gathered}
$$

The general form is given by Faà di Bruno's formula

$$
h_{n}=\sum_{k=1}^{n} f_{k} \mathrm{~B}_{n, k}\left(g_{1}, g_{2}, \ldots, g_{n-k+1}\right)
$$

Here $\mathrm{B}_{n, k}$ denotes a homogeneous polynomial of degree $k$ and weight $n$ in the $g_{m}$, called the Bell polynomial (see [2]). Another notation is

$$
\begin{equation*}
\mathrm{B}_{n}\left(g_{1}, g_{2}, \ldots, g_{n}\right)=\sum_{k=1}^{n} \mathrm{~B}_{n, k}\left(g_{1}, g_{2}, \ldots, g_{n-k+1}\right) \tag{1}
\end{equation*}
$$

which is the complete Bell polynomial.
They have the explicit form

$$
\begin{equation*}
\mathrm{B}_{n, k}\left(g_{1}, \ldots, g_{n-k+1}\right)=\sum \frac{n!}{j_{1}!\ldots j_{n-k+1}!}\left(\frac{g_{1}}{1!}\right)^{j_{1}} \ldots\left(\frac{g_{n-k+1}}{(n-k+1)!}\right)^{j_{n-k+1}} \tag{2}
\end{equation*}
$$

here the sum runs over the set of all partitions of $n$ with $k$ parts, or in other words, over all solutions in non-negative integers $j_{m}$ of the equations

$$
\sum_{m=1}^{n-k+1} j_{m}=k, \sum_{m=1}^{n-k+1} m j_{m}=n
$$

Another important formula (We will use it to obtain the general term of an asymptotic series.) is the generating function of the complete Bell polynomials given by

$$
\begin{equation*}
\mathcal{B}(z)=\exp \left(\sum_{m \geq 1} \frac{g_{m} z^{m}}{m!}\right)=1+\sum_{n \geq 1} \mathrm{~B}_{n}\left(g_{1}, g_{2}, \ldots, g_{n}\right) \frac{z^{n}}{n!} . \tag{3}
\end{equation*}
$$

There is also a more general application in the study of power series, for example let

$$
f(z)=\sum_{m \geq 0} a_{m} \frac{z^{m}}{m!}, g(z)=\sum_{m \geq 0} b_{n} \frac{z^{n}}{n!}
$$

then

$$
f(g(z))=\sum_{n \geq 0} \frac{\sum_{k=1}^{n} a_{k} B_{n, k}\left(b_{1}, b_{2}, \ldots, b_{n-k+1}\right) z^{n}}{n!}
$$

2. Stirling's series. Before introducing the new Gamma function approximation, let us first review the proof of the well-known Stirling series (see [1]).

Theorem 1 Let $z \geq 1$, then

$$
\begin{equation*}
\ln (\Gamma(z))=\left(z-\frac{1}{2}\right) \ln (z)-z+\frac{1}{2} \ln (2 \pi)+\sum_{n \geq 1} \frac{B_{2 n}}{2 n(2 n-1) z^{2 n-1}} \tag{4}
\end{equation*}
$$

The series is divergent for all values of $z$, but the partial sums can be made an arbitrarily good approximation for large enough $z$. Here the $B_{n}$ denote the Bernoulli numbers (see [1]).

Proof. In the proof we will apply the Euler-Maclaurin Summation Formula, the Trigamma function and the Wallis product.

The Trigamma function is defined by

$$
\psi_{1}(z)=\frac{d^{2} \ln (\Gamma(z))}{d z^{2}}=\sum_{n \geq 0} \frac{1}{(z+n)^{2}}
$$

We apply the Euler-Maclaurin Formula to find an asymptotic approximation to the function

$$
\int_{0}^{\infty} \frac{d x}{(z+x)^{2}}=\frac{1}{z}
$$

We get the expression

$$
\frac{1}{z}=\int_{0}^{\infty} \frac{d x}{(z+x)^{2}}=\frac{1}{2 z^{2}}+\sum_{k \geq 0} \frac{1}{(z+1+k)^{2}}-\sum_{n \geq 1} \frac{B_{2 n}}{z^{2 n+1}}
$$

By plugging the Trigamma function into this formula and solving for it we get

$$
\psi_{1}(z+1)=\frac{1}{z}-\frac{1}{2 z^{2}}+\sum_{n \geq 1} \frac{B_{2 n}}{z^{2 n+1}}
$$

Repeated integration gives an asymptotic series for the Gamma function which, however, contains an integration constant $\omega$.

$$
\begin{equation*}
\ln (\Gamma(z+1))=\omega+\left(z+\frac{1}{2}\right) \ln (z)-z+\sum_{n \geq 1} \frac{B_{2 n}}{2 n(2 n-1) z^{2 n-1}} \tag{5}
\end{equation*}
$$

In the limit case, this gives the following asymptotic approximation for $\Gamma(z+1)$.

$$
\Gamma(z+1) \approx z^{z+\frac{1}{2}} e^{-z} e^{\omega}, \text { as } z \rightarrow \infty
$$

To evaluate the constant we apply the Wallis product for $\pi$,

$$
\frac{\pi}{2}=\prod_{n \geq 1} \frac{(2 n)^{2}}{(2 n-1)(2 n+1)}
$$

Some manipulation gives the

$$
\begin{equation*}
\sqrt{\frac{\pi}{2}}=\lim _{n \rightarrow \infty} \frac{(2 n)!!}{\sqrt{2 n}(2 n-1)!!}=\lim _{n \rightarrow \infty} \frac{2^{2 n}(n!)^{2}}{\sqrt{2 n}(2 n)!} \tag{6}
\end{equation*}
$$

expression. We also have

$$
\Gamma(n+1)=n!\approx n^{n+\frac{1}{2}} e^{-n} e^{\omega}
$$

if $n$ is an integer. Inserting it into (6) we arrive at

$$
\lim _{n \rightarrow \infty} \frac{2^{2 n}\left(n^{2 n+1} e^{-2 n} e^{2 \omega}\right)}{(2 n)^{2 n+\frac{1}{2}} e^{-2 n} e^{\omega}} \frac{1}{\sqrt{2 n}}=\sqrt{\frac{\pi}{2}} \Leftrightarrow \omega \rightarrow \frac{1}{2} \ln (2 \pi) .
$$

Replacing $\omega$ in (5) by $\frac{1}{2} \ln (2 \pi)$ completes the proof.
3. The new asymptotic series. In this section we introduce a new series approximation to the Gamma function for large values of $z$.

Theorem 2 Let $z \geq 1$, then the following asymptotic series is valid

$$
\begin{equation*}
\Gamma(z)=\sqrt{\frac{2 \pi}{z}} e^{-z}\left(z+\sum_{n \geq 1} \frac{b_{n}}{z^{n-1}}\right)^{z} \tag{7}
\end{equation*}
$$

as $z \rightarrow \infty$. The $b_{n}$ coefficients are given by

$$
\begin{equation*}
b_{n}=\mathrm{B}_{n}\left(0, B_{2}, \ldots,(n-2)!B_{n}\right) \frac{1}{n!} . \tag{8}
\end{equation*}
$$

Proof. First we rewrite Stirling's series in the form

$$
\frac{1}{z} \ln \left(\Gamma(z)\left(\frac{e}{z}\right)^{z} \sqrt{\frac{z}{2 \pi}}\right)=\sum_{n \geq 1} \frac{B_{2 n}}{2 n(2 n-1) z^{2 n}}
$$

Taking the exponential of each side gives

$$
\begin{equation*}
\frac{e}{z} \sqrt[z]{\Gamma(z) \sqrt{\frac{z}{2 \pi}}}=\exp \left(\sum_{n \geq 1} \frac{B_{2 n}}{2 n(2 n-1) z^{2 n}}\right) \tag{9}
\end{equation*}
$$

To expand the right side of the equation to a series we can apply the generating function of the complete Bell polynomial (3) that we mentioned above. Here the variable is $z^{-1}$ instead of $z$. If we use the same notations the $g_{n}$ coefficients take the form

$$
g_{1}=0 \text { and } g_{n}=\frac{n!B_{n}}{n(n-1)}=(n-2)!B_{n}, \quad \text { if } n \geq 2
$$

Since $B_{3}=B_{5}=B_{7}=\ldots=0$. This means that our expression (9) becomes

$$
\frac{e}{z} \sqrt[z]{\Gamma(z) \sqrt{\frac{z}{2 \pi}}}=1+\sum_{n \geq 1} \mathrm{~B}_{n}\left(0, B_{2}, \ldots,(n-2)!B_{n}\right) \frac{1}{n!z^{n}}
$$

By solving for $\Gamma(z)$ we finally arrive at our asymptotic formula and it completes the proof.

The computation of the coefficients. From (1), (2) and (8) we can compute the numerical values of $b_{n}$. The Bernoulli numbers start

$$
B_{n}=1,-\frac{1}{2}, \frac{1}{6}, 0,-\frac{1}{30}, 0, \frac{1}{42}, 0,-\frac{1}{30}, \ldots(n \geq 0)
$$

So $b_{n}=0$ if $n$ is odd and the first few values of the even-indexed coefficients are

$$
\begin{aligned}
& b_{2}=\frac{1}{2} B_{2}=\frac{1}{12} \\
& b_{4}=\frac{1}{12} B_{4}+\frac{1}{8} B_{2}^{2}=\frac{1}{1440} \\
& b_{6}=\frac{1}{30} B_{6}+\frac{1}{24} B_{2} B_{4}+\frac{1}{48} B_{2}^{3}=\frac{239}{362880} \\
& b_{8}=\frac{1}{56} B_{8}+\frac{1}{60} B_{2} B_{6}+\frac{1}{288} B_{4}^{2}+\frac{1}{96} B_{4} B_{2}^{2}+\frac{1}{384} B_{2}^{4}=-\frac{46409}{87091200}
\end{aligned}
$$

Thus the expansion (7) starts

$$
\Gamma(z)=\sqrt{\frac{2 \pi}{z}} e^{-z}\left(z+\frac{1}{12 z}+\frac{1}{1440 z^{3}}+\frac{239}{362880 z^{5}}-\frac{46409}{87091200 z^{7}}+\ldots\right)^{z}
$$

| $b_{2}$ | 0.08333333333333333333333333333333333 |
| :--- | ---: |
| $b_{4}$ | 0.00069444444444444444444444444444444 |
| $b_{6}$ | 0.00065861992945326278659611992945326 |
| $b_{8}$ | -0.00053287817827748383303938859494415 |
| $b_{10}$ | 0.00079278588700608376534302460228386 |
| $b_{12}$ | -0.00184758189322033028400606295961969 |
| $b_{14}$ | 0.00625067824784941846328836824623616 |
| $b_{16}$ | -0.02901710246301150993444701506844402 |
| $b_{18}$ | 0.17718457242491308890302832366796470 |
| $b_{20}$ | -1.37747681703993534399676348903067470 |

Table with numerical values of $b_{n}$.

Peter Luschny showed that this asymptotic series can be expanded into a half-integer continued fraction formula

$$
\begin{equation*}
n!=\sqrt{2 \pi}\left(n+\frac{1}{2}\right)^{\left(n+\frac{1}{2}\right)} e^{-\left(n+\frac{1}{2}\right)}\left(\frac{\left(n+\frac{1}{2}\right)}{\left(n+\frac{1}{2}\right)+\frac{\frac{1}{24}}{\left(n+\frac{1}{2}\right)+\frac{\frac{3}{80}}{\ddots .}}}\right)^{\left(n+\frac{1}{2}\right)} \tag{10}
\end{equation*}
$$

This formula is particularly useful if we want to approximate the factorial function for both small and large values of $n$ (see [4]).
4. Numerical comparisons. Although we considered an asymptotical formula, i. e. a formula which is optimized for use with large values of $x$, it is for practical purposes also of interest to know the behaviour for small values of $x$. Therefore we will compare in this paragraph the numerical performance of some asymptotic formula to the Gamma function with our formula in the range $x \in[0.5$.. 50].

To this aim we introduce the relative error of an approximation $a(x)$ to $\Gamma(x)$. This can be defined as

$$
\begin{equation*}
\delta(x, a)=1-\frac{a(x)}{\Gamma(x)} \tag{11}
\end{equation*}
$$

We compare the following approximation formulas

$$
\begin{aligned}
& \sigma(x)=\left(\frac{x}{e}\right)^{x} \sqrt{\frac{2 \pi}{x}} \exp \left(\frac{1}{12 x}-\frac{1}{360 x^{3}}+\frac{1}{1260 x^{5}}-\ldots\right) \quad(\text { Stirling series) } \\
& \lambda(x)=\left(\frac{x}{e}\right)^{x} \sqrt{\frac{2 \pi}{x}}\left(1+\frac{1}{12 x}+\frac{1}{288 x^{2}}-\frac{139}{51840 x^{3}}-\ldots\right) \quad \text { (Laplace formula) } \\
& \rho(x)=\left(\frac{x}{e}\right)^{x} \sqrt{\frac{2 \pi}{x}} \sqrt[6]{\left(1+\frac{1}{2 x}+\frac{1}{8 x^{2}}+\frac{1}{240 x^{3}}-\ldots\right)} \quad \text { (Ramanujan formula), } \\
& \nu(x)=\left(\frac{x}{e}\right)^{x} \sqrt{\frac{2 \pi}{x}}\left(1+\frac{1}{12 x^{2}}+\frac{1}{1440 x^{4}}+\frac{239}{362880 x^{6}}-\ldots\right)^{x} \quad \text { (Nemes formula). }
\end{aligned}
$$

The second expression is sometimes incorrectly called Stirling's formula (see [6]).
The following graph showes the relative error of these formulas to the Gamma function. The plotted quantity is $|\ln (a(x) / \Gamma(x))| \approx|1-a(x) / \Gamma(x)|$. The first computed values are for $x=0.5$. Thick traces indicate $a(x)>\Gamma(x)$, thin ones $a(x)<\Gamma(x)$. Color codes for various families: blue - Stirling, green - Laplace, brown - Ramanujan, red - Nemes. The number after a name indicates the power of $(1 / x)$ in the last kept term in the expansion. The following curves overlap: Laplace_ 2 overlaps Stirling_1, Ramanujan_2 overlaps Nemes_ 2 for large $x$, Nemes_ 4 overlaps Nemes_ Closed, Nemes_6 overlaps Stirling_ 5. For Nemes_Closed formula, see the Section 5.


Figure 1: Relative errors of Gamma function approximations.
Conclusion. From the graph we see that the first and the third Laplace approximations outperform the corresponding Ramanujan formulas, but these expansions do not contain the same number of terms. Ramanujan_2 gives better approximation than Laplace_ 2 and Nemes_2, though the latter is nearly identical for larger values of $x$. The graph also shows that, for equal number of 'correction' terms, the Nemes formulae are always better than all the other (e.g., Nemes_ 4 is better than Stirling_ 3, Ramanujan_2 and Laplace_2). The Stirling and Nemes formulas are very useful, because they use only half of the powers of the variable contrary to the Laplace and Ramanujan series.

The behavior of the closed formula is very interesting, it gives approximately the same value as Nemes_ 4 and a better one than Stirling_ 3 , even though it contains only one 'correction' term.
5. An interesting closed approximation. We can get a good closed approximation to the Gamma function if we approximate the Laplace series by a much simpler function. Write up Laplace's formula in the form

$$
\Gamma(z)=\left(\frac{z}{e}\right)^{z} \sqrt{\frac{2 \pi}{z}}\left(1+\frac{1}{12 z}+\frac{1}{288 z^{2}}-\frac{139}{51840 z^{3}}-\frac{571}{2488320 z^{4}}+\frac{163879}{209018880 z^{5}}+\ldots\right)
$$

There exists a function which has a very similar series expansion

$$
\left(1+\frac{1}{15 z^{2}}\right)^{\frac{5}{4} z}=1+\frac{1}{12 z}+\frac{1}{288 z^{2}}-\frac{139}{51840 z^{3}}-\frac{571}{2488320 z^{4}}+\frac{16997}{149299200 z^{5}}+\ldots
$$

By exchanging the series for this function we get

$$
\begin{equation*}
\Gamma(z) \approx\left(\frac{z}{e}\right)^{z} \sqrt{\frac{2 \pi}{z}}\left(1+\frac{1}{15 z^{2}}\right)^{\frac{5}{4} z} \tag{12}
\end{equation*}
$$

which is valid for large values of $z$. This is Nemes_Closed in the comparison paragraph.
Acknowledgement. I want to thank Peter Luschny, Péter Simon and Stanislav Sýkora for discussing with me some of the points in this paper.

Comments. Peter Luschny gave another definition of the $b_{n}$ coefficients which is equivalent to the formula we obtained above. If we write our formula in the form

$$
\begin{equation*}
\Gamma(z)=\sqrt{\frac{2 \pi}{z}}\left(\frac{z}{e} \sum_{n \geq 0} \frac{b_{2 n}}{z^{2 n}}\right)^{z} \tag{13}
\end{equation*}
$$

the $b_{2 n}$ coefficients can be computed by

$$
\begin{equation*}
\sum_{\sum_{1}^{j} p_{i} \mu_{i}=2 n} \frac{B_{p_{1}}^{\mu_{1}}}{\mu_{1}!\left(p_{1}^{2}-p_{1}\right)^{\mu_{1}}} \frac{B_{p_{2}}^{\mu_{2}}}{\mu_{2}!\left(p_{2}^{2}-p_{2}\right)^{\mu_{2}} \ldots} \frac{B_{p_{j}}^{\mu_{j}}}{\mu_{j}!\left(p_{j}^{2}-p_{j}\right)^{\mu_{j}}} . \tag{14}
\end{equation*}
$$

Here the sum runs over all partitions of $2 n$ with even parts.
The error curves were generated by Stanislav Sýkora using Matlab. For more information about this please visit his website: www.ebyte.it

## References

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